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# Deep-sea Fe-Mn Crusts from the Northeast Atlantic Ocean: Composition and Resource Considerations

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## Deep-sea Fe-Mn Crusts from the Northeast Atlantic Ocean: Composition and Resource Considerations

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Eighteen deep-sea ferromanganese crusts (Fe-Mn crusts) from 10 seamounts in the northeast Atlantic were studied. Samples were recovered from water depths of  $\sim$ 1,200 to  $\sim$ 4,600 m from seamounts near Madeira, the Canary and Azores islands, and one sample from the western Mediterranean Sea.

The mineralogical and chemical compositions of the samples indicate that the crusts are typical continental margin, hydrogenetic Fe-Mn crusts. The Fe-Mn crusts exhibit a Co + Cu + Ni maximum of 0.96 wt%. Platinum-group element contents analyzed for five samples showed Pt contents from 153 to 512 ppb.

The resource potential of Fe-Mn crusts within and adjacent to the Portuguese Exclusive Economic Zone (EEZ) is evaluated to be comparable to that of crusts in the central Pacific, indicating that these Atlantic deposits may be an important future resource.

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Address correspondence to Susana Bolhão Muiños, Instituto Português do Mar e da Atmosfera, I.P., Divisão de Geologia e Georecursos Marinhos, Rua C-Aeroporto de Lisboa, 1749-077 Lisboa, Portugal. E-mail: susana.muinos@ipma.pt **Keywords** ferromanganese crusts, northeast Atlantic, Portuguese EEZ, resource considerations, seamounts

## Introduction

Manganese and ferromanganese oxide deposits in the oceans occur as nodules, crusts, and massive beds. These deposits have been classified as diagenetic, hydrogenetic, hydrothermal, and mixed-type deposits (Halbach 1986; Hein et al. 1997; Wen et al. 1997). Hydrogenetic crusts (Fe-Mn crusts) form by direct precipitation of colloidal hydrated metal oxides from the water column onto hard-rock substrates.

The first investigations of hydrogenetic Fe-Mn crusts on seamounts were carried out in the Pacific Ocean (Craig et al. 1982; Halbach et al. 1982, 1989b; Hein et al. 1988). The preconditions required for Fe-Mn crust formation, such as the occurrence of isolated volcanic edifices, strong currents that keep the edifices free of sediment, and an oxygen-minimum zone (OMZ) are also found in the Atlantic Ocean (Koschinsky et al. 1995). Indeed, previous results from NE Atlantic seamounts indicate widespread presence of Fe-Mn crusts of hydrogenetic origin (Koschinsky et al. 1995, 1996; Gaspar 2001; Muiños et al. 2002; Muiños 2005). Hydrogenetic precipitation is dependent on water-mass properties and is characterized by slow growth rates (< 10 mm/Ma) and generation of an extremely high specific-surface area, which promotes the enrichment of trace elements through scavenging by the major oxides (e.g., Hein et al. 1997). Seamounts act as obstructions to oceanic water-mass flow thereby creating seamount-generated currents of enhanced energy relative to flows away from the seamounts. These currents, which are strongest along the outer rim of the summit region of seamounts, promote the formation of thick crusts, enhanced turbulent mixing, and produce upwelling, leading to increased primary productivity and thus maintenance of the OMZ (summarized in Hein et al. 2000). Manganese oxides and associated trace metals are concentrated in the OMZ, which are then scavenged onto crusts under oxic conditions resulting from the turbulent mixing around seamounts.

Our study area in the northeast Atlantic (Figure 1) is influenced by the Mediterranean Outflow Water (MOW), which is characterized by relatively high salinity and temperature and low oxygen content compared to surrounding water masses. Because manganese is soluble under low-oxygen conditions, this water mass is a reservoir for  $Mn^{2+}$  in solution between 800 and 1,200 m water depth, which corresponds to the upper and lower cores of the MOW (Madelain 1970; Zenk 1970; Ambar and Howe 1979; Ambar et al. 1999, 2002). Fluctuations in the intensity of the OMZ and MOW may have influenced the composition of the Fe-Mn crusts (Koschinsky et al. 1996).

Hydrogenetic precipitation promotes the enrichment of crusts in potentially economically important trace metals such as Co, Ni, Te, rare earth elements (REEs), and platinum-group elements (PGEs), and thus there is a growing recognition of Fe-Mn crusts as potential metal resources. With growing markets for metals in Asia, as well as the rapid development of high-tech and green-tech applications, the demand for rare earth metals will increase dramatically in the near future (Hein et al. 2010).



Figure 1. Map showing bathymetry, the location of the sample sites, and the limits of the Portuguese EEZ. (Color figure available online.)

Significantly, to our knowledge, there is a lack of technology needed for detailed exploration and extraction of Fe-Mn crusts, which may be slowing development. However, much proprietary engineering research has been undertaken in a number of countries. Environmental studies are also in their infancy and should be addressed as stipulated in the United Nations Convention on the Law of the Sea and International Seabed Authority regulations. Accompanying the accelerating economic interest in Fe-Mn crusts, the International Seabed Authority passed regulations for exploration for Fe-Mn crusts during its 18th Session in July 2012.

Much work has been done on marine Fe-Mn deposits, but studies of Atlantic Ocean deposits are still scarce compared to those from the Pacific Ocean. The aims of this study are to determine the composition of Fe-Mn crusts from the northeast Atlantic, in particular those within the Portuguese Exclusive Economic Zone (EEZ), and to consider their resource potential.

### Material and Methods

Samples were collected from 10 seamounts in the northeast Atlantic during cruises Trident 86, TTR-11 (I.O.C. 2002), Meteor 51/1 (Hoernle and Scientific Party 2003), and Tore-Madeira (Merle 2006). Sampling occurred over a wide geographic range including seamounts in the Portuguese EEZ and near the Canary Islands. Sampling locations are distributed over a large depth range and the growth and composition of the Fe-Mn crusts were influenced by different chemical and oceanographic environments that are broadly representative of the study area (Figure 1 and Table 1).

The chemical and mineralogical data represent analyses of bulk samples. All crusts were analyzed by X-ray diffraction on a Philips diffractometer using Cu-K $\alpha$  radiation and carbon monochromator at the United States Geological Survey (USGS). The interpretation of the diffractograms and identification of mineralogical phases were also performed at the USGS using the program X'Pert High Score of Philips (PANalytical). We follow the nomenclature of Usui et al. (1989) for manganese minerals. The semi-quantitative determination of mineral content is based on

TADIC T. FO	cauon or samples					
Sample ID	Field ID	Latitude (N)*1	Longitude (W)* <sup>1</sup>	Water Depth (m)* <sup>1</sup>	Seamount	Geographic Area
3477- <b>B</b> .3.4	TTR11-353GR	35.3117	-14.8300	1853	Nameless	Madeira
3478- <b>B</b> .3.5	TTR11-354GR	35.3167	-14.8350	1839	Nameless	Madeira
3717-12	TM-D19	35.2767	-14.8500	2198	Nameless	Madeira
3521-6	M51/1-426DR	33.8683	-14.3680	1362	Seine	Madeira
3522-4	M51/1-428DR	34.4567	-15.5387	2946	Godzilla	Madeira
3718-1	TM-D20	34.4258	-14.4173	3756	Unicorn	Madeira
3513-13	M51/1-414DR	36.9700	-14.7475	4594	MTR-Josephine	Madeira
3513-14	M51/1-414DR	36.9700	-14.7475	4594	MTR-Josephine	Madeira
3513-16	M51/1-414DR	36.9700	-14.7475	4594	MTR-Josephine	Madeira
3708-1	TM-D3B	39.7217	-13.7152	4140	Tore	Madeira
3709-1	TM-D5	38.8442	-13.0147	2803	Tore	Madeira
3710-1	TM-D6B	39.2158	-12.8464	4245	Tore	Madeira
3711-2	TM-D9	39.2775	-12.1486	3110	Tore	Madeira
3525-9	M51/1-448DR	31.3216	-13.8009	3043	Dacia	Canaries
3533-7	M51/1-457DR	31.6267	-12.9382	1602	Annika	Canaries
3534-14.1	M51/1-458DR	31.7083	-12.8156	2128	Annika	Canaries
3536-3	M51/1-462DR	35.8450	-4.2125	1221	Ibn Batouta	Mediterranean
3862	TR86-8D	36.5667	-26.4833	2575	Azores	Azores
*1Except foi TM-Tore-Mac	r samples from the T leira and TR-Trident	TR-11 cruise that we 86) correspond to int	re collected by TV-Gra ermediate values from o	ib, all the other coordinat on- and off-bottom dredge	tes and water depths () closed of the closed	M51/1- Meteor 51/1, s Madeira-Tore Rise.

Table 1. Location of samples

the relative intensity of the peaks and previously determined weighing factors (Cook et al. 1975; Hein et al. 1988).

Major elements (Fe, Mn, Si, Al, Ca, Mg, Na, K, Ti, P) were analyzed by fuseddisk X-ray fluorescence; S, Ba, Cr, Cu, Li, Ni, Sr, V, Zn, and Zr by 4-acid digestion and inductively coupled plasma-optical emission spectrometry (ICP-OES); Ag, As, Be, Bi, Cd, Co, Ga, Ge, Hf, In, Mo, Nb, Pb, Rb, Sb, Sc, Sn, Ta, Tl, W, and Cs were analyzed by 4-acid digestion and ICP-mass spectrometry (MS); Th, U, Y, and REEs were analyzed by lithium metaborate fusion and ICP-MS; Se and Te by 4-acid digestion, hydride-generation, and atomic absorption spectrometry (AAS), Hg by cold vapor AAS, and Cl<sup>-</sup> was analyzed by the specific-ion electrode method. Based on duplicate analyses of 10% of the samples, precision was better than 5% for S, As, Ba, Be, Bi, Cd, Co, Cr, Cu, Ga, Ge, In, Li, Mo, Ni, Pb, Rb, Sc, Se, Sr, Te, Th, Tl, U, V, Zn, REEs, Cl<sup>-</sup>, and Cs and better than 10% for Sb, Sn, W, and Hg. For a few elements, precision varies widely and data should be used with that in mind: Ag (10–33%), Hf (11–28%), Nb (15–24%), Ta (13–24%), and Zr (5–16%). Five samples were also analyzed for PGE and Au contents by fire assay and ICP-MS. Analytical accuracy was calculated using international standards AMIS0056 and HGMNEW and is better than 5% for Os and Ru, better than 10% for Ir, Pd, and Pt, 13% for Au, and varies from 5-17% for Rh.

Q-Mode factor analyses used the Varimax method (Klovan and Imbrie 1971). All communalities are >0.94 and values between -0.1575 and 0.1575 were not considered because they are below the level of statistical significance assuming a multi-Gaussian distribution.

The extent of the area covered by seamounts was determined using ArcGIS<sup>®</sup> and ETOPO bathymetry (Amante and Eakins 2009; http://www.ngdc.noaa.gov/mgg/global/global.html).

## Mineralogy and Chemical Composition of Crusts

All Fe-Mn crust samples are composed predominately of  $\delta$ -MnO<sub>2</sub> (vernadite), the mineral most characteristic of hydrogenetic Fe-Mn deposits found globally. The mineral  $\delta$ -MnO<sub>2</sub> is epitaxially intergrown with X-ray amorphous iron oxyhydroxide (δ-FeO(OH)-feroxyhyte; Burns and Burns 1977; Varentsov et al. 1991; Hein et al. 2000). Detrital minerals, such as quartz and feldspar, and biogenic and diagenetic minerals, such as calcite and carbonate fluorapatite (CFA) are present in minor to moderate amounts (Table 2). In addition, two samples contain minor amounts of 10 A manganate (probably todorokite), which may reflect a lower oxidation potential of seawater caused by increased biological productivity, as suggested by Hein et al. (2000) for some occurrences in the Pacific Ocean, or may indicate a minor hydrothermal contribution. Minor amounts of goethite are also present in the majority of the samples and can reflect a number of different processes, including: (1) increased inputs of Fe from continental sources (Bruland et al. 2001); (2) too much Fe present for the vernadite structure to accommodate (De Carlo 1991, and references therein); (3) the enhanced supply of Fe from the dissolution of calcareous tests for crusts below the calcite compensation depth (CCD), as suggested by von Stackelberg et al. (1984) for goethite in some layers of a deep-water Fe-Mn crust (4,830m) collected in the Clarion-Clipperton nodule belt; and (4) as suggested by Hein et al. (2000), goethite is found only in the older parts of 5% of 640 crust samples analyzed from Pacific Ocean sites and may result from the maturation of X-ray amorphous Fe oxyhydroxide. Most

Sample	δ-MnO <sub>2</sub> (%)	Others* <sup>2</sup>
3477-B.3.4	95	feldspar; todorokite; quartz; goethite
3478-B.3.5	93	goethite; calcite; plagioclase; quartz
3513-13	94	goethite; plagioclase; quartz
3513-14	94	goethite; quartz; plagioclase
3513-16	85	quartz; goethite; feldspar; kaolinite
3521-6	92	CFA; low Mg-calcite; goethite
3522-4	99	calcite
3708-1	94	quartz; goethite
3709-1	91	quartz; goethite; plagioclase
3710-1	95	quartz; goethite
3711-2	87	calcite; quartz; goethite
3717-12	97	goethite; plagioclase; quartz
3718-1	87	quartz; goethite; aragonite; plagioclase
3525-9	99	calcite; quartz
3533-7	85	CFA; goethite
3534-14.1	93	goethite; feldspar; quartz; calcite
3536-3	71	calcite; Mg-calcite; todorokite; goethite; quartz
3862	96	goethite

**Table 2.** X-ray Diffraction Mineralogy of Fe-Mn samples

\*<sup>2</sup>In approximate decreasing abundance; CFA means carbonate fluorapatite.

Atlantic crusts in our study do not originate from below or near the CCD and their Fe concentrations are not greater than those of many crusts without goethite, so that explanations (1) and (4) may be most likely for the occurrence of goethite in the Atlantic samples.

The Fe-Mn crust samples have Fe contents from 12.5 to 23.2 wt% with an average of 17.9 wt% and Mn contents from 9.3 to 17.0 wt% with a mean of 13.7 wt% (Table 3). These concentrations result in Fe/Mn ratios of 0.88 to 1.96 with a mean of 1.33, which is in the range typical of continental margin hydrogenetic Fe-Mn crusts (Hein et al. 2000). Silicon and Al concentrations range from 1.21 to 13.0 wt% and 0.73 to 4.66 wt% with average values of 4.45 and 2.04 wt%, respectively; Si/Al ratios vary between 0.83 and 5.32, with an average value of 2.05. The Si/Al ratios of the volcanic rocks in the area (Geldmacher et al. 2006; Merle et al. 2006) cluster around 2.8. Sample 3708-1 from Unicorn Seamount has a Si/Al ratio of 5.32 indicating the presence of biogenic silica or eolian quartz. An explanation for the low Si/Al ratios is more difficult but preferential incorporation of feldspar over quartz or sorption of Al hydroxide as suggested by Koschinsky and Halbach (1995) could explain those ratios. Calcium concentrations range from 0.99 to 12.7 wt% with a mean value of 3.23 wt%, and P varies from 0.21 to 4.28 wt%, averaging 0.65 wt%.

Cobalt, Cu, and Ni show minimum concentrations of 931, 315, and 1,240 ppm and maximum concentrations of 5,440, 2,050, and 3,980 ppm, respectively. Cobalt has a mean concentration of 3,408 ppm, copper averages 872 ppm, and nickel 2,197 ppm. The Co + Cu + Ni maximum is 0.96 wt%, with a mean of 0.64 wt%. These values and the Fe/Mn ratios are consistent with a predominantly hydrogenetic origin of continental margin Fe-Mn crusts. Some compositional data points fall outside

Table 3. Chemical composition

3521-6		DUP-3521-6	3522-4	3525-9	3536-3	3708-1	3709-1	3710-1	3711-2	3718-1
wt% 17.8 N.A. 20	N.A. 20	20	.3	22.9	12.5	17.8	16.9	20.8	18.0	17.9
13.5 N.A. 16	N.A. 16	16		14.3	14.2	13.8	15.3	10.6	11.9	12.1
1.21 N.A. 2.2	N.A. 2.2	2.2	0	2.34	3.65	5.94	6.78	4.96	3.87	9.77
1.47 N.A. 1.4	N.A. 1.4	1.4	9	1.33	2.62	3.05	2.07	3.05	1.76	1.84
6.05 N.A. 2.5	N.A. 2.5	2.5	4	2.16	8.58	1.50	1.72	1.40	5.06	1.49
1.92 N.A. 1.1	N.A. 1.1	1.1	8	1.09	4.06	1.49	1.21	1.28	1.08	0.98
0.91 N.A. 1.0	N.A. 1.0	1.0	5	1.00	0.53	1.12	1.31	0.87	1.00	1.26
0.25 N.A. 0.3	N.A. 0.3	0.3	0	0.29	0.49	0.48	0.51	0.44	0.43	0.60
1.10 N.A. 0.9	N.A. 0.9	0.9	1	0.98	0.10	0.80	0.61	0.93	0.88	0.50
1.46 N.A. 0.4	N.A. 0.42	0.4	0	0.46	0.21	0.41	0.33	0.46	0.38	0.29
0.39 0.40 0.32	0.40 0.32	0.32		0.28	0.23	0.26	0.24	0.21	0.25	0.20
1.3 – 1.3	- 1.3	1.3		1.6	0.9	1.3	1.1	2.0	1.5	1.5
0.83 – 1.51	- 1.51	1.51		1.75	1.39	1.95	3.27	1.62	2.20	5.32
ppm 0.29 0.10 0.34	0.10 0.34	0.34		0.23	0.04	0.12	0.12	0.11	0.11	0.24
455 416 317	416 317	317		373	328	250	233	298	272	244
884 904 1270	904 1270	127(	0	1450	953	779	1240	740	882	1540
9.5 9.0 11	9.0 11	11		12	3.7	9.3	8.6	8.3	7.8	8.7
33 31 31	31 31	31		25	6.5	22	24	24	22	16
4.0 3.7 4.6	3.7 4.6	4.6		4.0	2.8	7.4	3.8	5.3	2.9	2.4
5440 5230 400	5230 400	400	0	3380	931	3250	3440	3180	3260	2510
101 109 21	109 21	21		78	61	49	121	54	35	30
426 422 784	422 784	784	_	592	317	1420	1040	786	368	1410
8.6 8.1 9.6	8.1 9.6	9.6		13	7.8	14	15	13	9.8	14
0.8 0.8 0.	0.8 0.	0	Ľ	0.9	0.2	0.6	0.5	0.6	0.7	0.4
8.4 3.7 8.	3.7 8.	<u></u> .	5	3.9	0.36	3.3	1.7	3.3	1.6	7.1
0.12 0.11 0.1	0.11 0.1	0.1	0	0.12	0.05	0.12	0.14	0.20	0.10	0.13

11	446	43	1440	1000	19	47	9.1	<0.2	4.4	916	0.63	32	49	60	6.6	764	53	622	128	141	1280	32.3	112	25.8	5.68	22.6	3.62
11	299	31	1240	1320	15	32	14	0.4	4.9	1070	0.60	42	99	65	12	793	64	435	212	327	1360	80.1	298	70.0	16.3	71.2	11.0
22	228	41	1290	1120	15	43	20	0.6	10	719	1.2	46	75	75	11	795	48	473	180	229	1540	57.8	214	53.2	13.1	58.5	9.34
22	387	41	2650	1230	14	45	11	0.3	6.3	1040	0.42	43	46	140	10	(99)	80	548	57	211	1370	51.7	188	43.6	10.1	42.2	6.62
49	308	34	2570	1110	14	38	15	0.5	4.8	849	0.85	34	40	116	11	737	99	594	173	196	1100	49.1	186	46.6	11.0	50.0	8.27
117	184	5.2	2790	519	18	26	12	<0.2	1.7	557	O.05	11	24	38	4.7	599	36	566	16	53.8	250	14.2	55.3	15.8	3.79	17.6	2.96
9	464	50	1400	1740	6.6	57	16	0.5	5.3	1320	1.8	42	44	49	14	1030	108	560	72	371	1570	84.1	316	71.6	17.1	78.4	12.3
8	444	58	1830	1880	6.1	46	13	0.4	5.4	1290	1.8	47	37	142	14	962	109	557	137	294	1270	60.9	251	57.2	13.9	63.3	9.99
18	314	36	2420	1930	4.9	39	18	<0.2	6.7	1400	1.2	67	54	112	15	766	50	475	240	343	1080	74.8	286	66.8	16.2	74.1	11.6
18	328	68	2350	1860	5.4	46	20	$<\!0.2$	7.6	1380	2.3	71	54	118	15	914	64	463	214	356	1120	77.8	297	70.1	16.7	77.7	12.1
mqq																											
Li	Mo	Nb	ïŻ	Pb	Rb	Sb	Sc	Se	$\operatorname{Sn}$	Sr	Та	Te	Th	Π	D	>	W	Zn	Zr	La	Ce	$\mathbf{Pr}$	Νd	Sm	Eu	Gd	Tb

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Table 3. Continued

		3521-6	DUP-3521-6	3522-4	3525-9	3536-3	3708-1	3709-1	3710-1	3711-2	3718-1
Dy		65.0	62.4	50.8	63.2	15.9	43.0	33.2	49.1	55.3	17.1
Y		340	333	184	230	9.77	164	103	182	186	46.1
Ho		14.9	14.3	11.1	13.5	3.50	9.54	7.03	10.7	11.6	3.35
$\mathrm{Er}$		42.1	41.3	29.7	37.6	10.1	26.8	19.2	30.0	31.4	9.45
Tm		5.90	5.81	4.29	5.21	1.48	3.96	2.74	4.50	4.30	1.38
$\mathbf{Y}\mathbf{b}$		35.6	34.7	26.1	32.1	8.9	24.6	17.3	28.0	25.6	8.8
Lu		5.62	5.37	3.93	4.73	1.34	3.87	2.60	4.32	3.75	1.37
Hg	qdd	14	14	13	٢	7	10	8	25	11	10
Cl-	mqq	>5000	>5000	>5000	>5000	3120	>5000	>5000	>5000	>5000	>5000
		3862	3477-B.3.4	3478-B.3.5	3513-13	DUP-3513-13	3513-14	3513-16	3533-7	3534-14.1	3717-12
Fe	wt%	23.2	17.6	16.6	16.6	N.A.	15.7	12.6	14.0	21.4	20.3
Mn		13.3	17.0	16.3	13.4	N.A.	13.2	9.3	13.5	13.2	15.5
Si		1.57	3.09	3.36	4.47	N.A.	5.94	13.0	1.30	2.48	2.32
$\mathbf{AI}$		1.29	1.74	1.73	2.45	N.A.	2.65	4.66	0.73	1.49	1.23
Ca		2.04	1.92	3.34	1.40	N.A.	1.28	0.99	12.7	2.17	1.87
Mg		1.24	1.48	1.57	1.32	N.A.	1.36	1.83	1.13	1.32	1.10
Na		1.13	1.25	1.16	1.22	N.A.	1.13	1.27	0.84	0.98	1.14
K		0.24	0.51	0.49	0.45	N.A.	0.68	1.59	0.37	0.31	0.46
Ti		2.37	0.67	0.57	0.83	N.A.	0.52	0.74	0.29	1.01	0.32
Р		0.51	0.38	0.41	0.39	N.A.	0.31	0.24	4.28	0.51	0.31
S		0.29	0.30	0.26	0.24	0.23	0.18	0.14	0.31	0.27	0.26
Fe/Mn		1.7	1.0	1.0	1.2	I	1.2	1.4	1.0	1.6	1.3
Si/Al		1.21	1.77	1.95	1.82	I	2.24	2.79	1.79	1.66	1.88
$\operatorname{Ag}$	uıdd	0.44	0.34	0.14	0.12	0.15	0.18	0.23	0.05	0.12	0.04

346	1550	11	25	2.7	1970	111	449	13	0.6	1.2	0.16	L	548	11	2210	1060	11	56	11	<0.2	4.9	1170	0.12	34	45	143	8.3
419	991	12	24	3.4	4330	09	454	10	0.7	3.0	0.09	11	350	50	1590	1650	6.5	47	17	0.4	5.0	1090	1.2	58	52	73	14
279	2180	7.3	15	2.5	2640	49	315	8.0	0.4	0.86	0.06	13	278	11	1990	1330	8.3	37	7.4	$<\!0.2$	2.1	1470	0.38	32	17	85	9.2
129	769	5.5	14	3.8	1650	27	2010	19	0.3	7.3	0.31	63	164	58	2590	539	34	27	17	0.3	3.4	484	1.8	33	50	61	5.1
207	1170	8.1	18	5.0	2020	67	2050	19	0.6	5.5	0.29	43	360	43	2590	863	20	34	13	<0.2	3.7	745	0.99	25	69	120	7.6
213	958	7.1	18	4.6	3800	33	1130	14	0.7	3.1	0.19	26	289	41	1990	953	12	32	14	0.3	3.8	1060	1.1	41	59	123	9.3
214	955	7.2	18	4.7	3860	31	1140	14	0.7	2.4	0.19	27	289	29	2010	943	12	32	14	0.3	3.8	1090	0.84	40	61	122	9.6
278	1360	8.0	23	4.0	4060	50	696	16	0.5	5.8	0.11	33	341	41	3980	1720	12	48	8.6	<0.2	5.5	1030	0.60	56	37	169	8.6
316	1320	8.8	30	3.7	4930	51	970	16	0.5	12	0.14	26	379	68	3670	2100	12	55	10	< 0.2	7.2	1130	1.2	99	39	195	9.7
439	1450	12	15	4.6	4270	72	382	14	1.0	6.5	0.10	11	371	74	1330	991	3.2	47	21	0.3	8.1	1340	4.1	49	43	73	19
bpm																											
$\mathbf{As}$	Ba	Be	Bi	Cd	Co	Cr	Cu	Ga	Ge	Ηf	In	Li	Мо	Nb	ïZ	Pb	$\mathbf{Rb}$	$\mathbf{Sb}$	Sc	Se	$\operatorname{Sn}$	$\mathbf{Sr}$	Та	Te	Th	Π	Ŋ

(Continued)

Table 3. Continued

	3867	3477-B 3 4	3478-R 3 5	3513-13	DIIP-3513-13	3513-14	3513-16	3533-7	3534-14 1	3717-17
	1000			CT CT CC			01 0100			71 11/0
^	1130	848	<i>T9T</i>	662	644	612	270	861	974	904
W	54	115	92	57	57	64	25	84	72	121
Zn	650	724	755	468	475	604	547	576	538	652
Zr	75	151	88	122	178	204	104	88	139	55
La	390	190	174	282	270	189	120	181	324	265
Ce	1460	1210	1100	1450	1400	1290	972	662	1260	1400
$\Pr$	90.8	40.8	39.2	70.3	67.2	52.8	30.9	36.8	71.7	61.8
Nd	339	153	145	254	244	191	113	142	269	229
Sm	78.9	34.4	33.4	61.6	59.2	48.3	27.9	31.9	62.7	52.4
Eu	18.8	8.10	7.85	14.5	13.8	11.2	6.50	7.38	14.9	12.4
Gd	85.0	35.0	34.4	62.0	59.5	46.6	27.0	36.1	68.5	51.1
Tb	13.6	5.61	5.30	9.87	9.55	7.37	4.32	5.62	10.8	7.87
Dy	69.8	29.0	26.8	49.3	47.5	36.5	21.6	30.7	55.8	38.4
Y	269	110	100	164	158	120	79.9	159	217	111
Ho	15.4	6.29	5.85	10.4	9.92	7.49	4.52	7.12	12.4	8.03
Er	41.8	17.6	16.1	27.9	27.1	20.8	11.9	20.0	34.0	21.4
Tm	5.82	2.48	2.26	4.03	3.80	2.92	1.78	2.84	4.80	2.97
Yb	35.0	15.5	14.0	23.8	23.6	18.4	11.1	17.3	29.3	18.3
Lu	5.19	2.48	2.26	3.76	3.54	2.75	1.68	2.67	4.42	2.81
Hg pr	b 292	L	8	8	6	9	8	9	6	8
Cl <sup>-</sup> pp	m >500	0 >5000	>5000	>5000	>5000	>5000	>5000	3640	>5000	>5000

All Cs concentrations <5 ppm.

50



**Figure 2.** Ternary diagram. All data plotted as wt %. The fields are: (A) diagenetic, (B) hydrogenetic, and (C) hydrothermal. Ternary plot software from (www.crog.org/cedric/dplot).

the typical open-ocean hydrogenetic field (Figure 2), but do fall within the field typical for Pacific continental margin and small ocean basin Fe-Mn crusts. These continental-margin-type crusts show lower total Co + Cu + Ni and higher Fe, Si, Al, and Cr contents than open-ocean crusts, reflecting the higher terrigenous input per unit area for the Atlantic (Koschinsky et al. 1995; Hein et al. 2000).

Tellurium is also an important element because of its potential economic value in the solar cell industry (Hein et al. 2010). In our samples, Te ranges from 11 to 71 ppm with an average value of 43 ppm. Hein et al. (2003) reported mean concentrations of Te for Pacific and Atlantic hydrogenetic Fe-Mn crusts of 50 ppm, which is consistent with our data. Vanadium, W, Zr, and Th also have notably high concentrations in our samples (Table 3).

Differences in crust composition based on subregions of our study area (Mediterranean, Canaries, Tore, Maderia-Tore Rise (MTR), and Azores) are presented in Figure 3 using data averaged from Table 3 for each subregion (see Table 1). It should be stressed that for Mediterranean and Azores subregions, only one sample is available, which should be taken in to consideration when comparing to the average values from the remaining subregions. We also distinguish between the different MTR seamounts: MTR-Josephine and South MTR seamounts (which includes Nameless, Unicorn, Seine and Godzilla seamounts). South MTR seamounts is the subregion with the highest average values of Co + Ni + Cu, Co, Ni, Zn, Mo, Tl, W, Te, and Bi. The Mediterranean subregion is systematically lower for all elements, with the exceptions of Ni and Zn. The Canaries subregion shows the highest P contents, which will be discussed in the section on phosphatization, and the Azores subregion shows the higher concentrations of Ti (2.37 wt%), which may include a significant detrital source as well as the typical adsorption of Ti from seawater.

Changes in the geochemical composition of Fe-Mn crusts with water depth, such as increased Cu, Ni, Si, and Al concentrations with increasing water depth have been identified previously (e.g., Cronan 1977, 1997; Mangini et al. 1987; Hein et al. 1997). Increased Ni contents with water depth are not found in our sample set (Figure 4).



Figure 3. Trace-metal concentrations in different subregions of the study area. Data correspond to averages of analyses in Table 3 (See also Table 1).

There is an apparent increase of Cu content with water depth, which might be caused by the release of Cu during dissolution of carbonate material close to the CCD (Halbach et al. 1979 and references therein; Cronan 1997; Verlaan et al. 2004). Planktonic organisms extract metals from surface waters for metabolic processes and through scavenging. Dissolution of the carbonate tests and oxidation of the organic matter then release these metals at depth. The CCD in the North Atlantic lies at about 4,700-4,800 m (van Andel 1975; Broecker and Peng 1982). Our most Cu-enriched samples are from  $\sim$ 4,600 m water depth. Cronan (1997) found that for Pacific manganese nodules, the highest Cu contents are not present in the deepest water nodules below the CCD, but in those located slightly above the CCD at around 5,000 m water depth. The increased deep-water Cu concentrations may thus be the result of a reduction in sedimentation rate associated with the loss of carbonate tests by dissolution near the CCD. That loss would lead to an increase in the concentration of organic phases in the sediments, whose decay would promote diagenetic reactions that enrich Cu and increase Al and Si contents (Cronan 1997). Our dataset supports this and shows similar increases with water depth for Cu, Al, and Si, and an opposite trend for Ca (Figure 4). Dissolution of carbonate tests may influence geochemistry



Figure 4. Distribution of selected elements with water depth. Dashed lines correspond to hand-drawn trends, others are regression lines; dashed box in Co illustrates the group of samples not considered for regression lines.

within the calcite lysocline at depths from about 2,600 to 3,000 m as indicated by a significant change in the slope of the Ca pattern and a change from increasing Fe/Mn to decreasing Fe/Mn with depth (Figure 4).

#### **Rare Earth Elements**

Chondrite-normalized REE patterns (Figure 5a) show enrichment of lightREEs (LREEs) relative to heavyREEs (HREEs), a strong positive Ce anomaly, and a small positive Gd anomaly, which are typical of Fe-Mn crusts of hydrogenetic origin. Shale-normalized patterns (Figure 5b) are also characteristic of hydrogenetic Fe-Mn crusts in that they show pronounced middleREEs (MREEs) enrichment (Nath et al. 1992 and references therein). No direct relationship with water depth or geographic location is discernable, with the exception of crust 3536-3 from the Mediterranean Sea in which REEs concentrations are much lower. All samples are highly enriched in REEs relative to seawater and the Earth's crust, with up to 0.29 wt.% total REEs, crust 3536-3 being the exception with only 475 ppm total REEs. This sample grew in the Mediterranean Sea and precipitation of the oxides was governed by waters in the Alboran Basin mixed with Atlantic waters that entered through the Strait of Gibraltar. The waters from the Mediterranean are characterized by low oxygen contents and a low redox potential. Also, this sample and sample 3533-7 show less pronounced Ce anomalies. The presence of positive Ce anomalies is caused by scavenging of Ce from seawater by hydrous Fe-Mn oxides (Goldberg et al. 1963; Elderfield et al. 1981) and its preferential retention relative to the other REEs in the oxide phase through surface oxidation (Bau et al. 1996). According to Kuhn et al. (1998), and given that Ce (III) oxidation is characterized by slow reaction kinetics (Sholkovitz and Schneider 1991; Moffet 1994), the size of the Ce anomaly will depend on the duration that the Fe-Mn precipitates are in contact with seawater. De Carlo (1991) suggested that variations in the REEs abundances and the extent of fractionation between LREEs and HREEs primarily reflect changes in the mineralogical composition of the crusts. De Carlo (1991) also pointed out that care must be taken when



**Figure 5.** (a) Chondrite-normalized REE patterns. Normalization values from Anders and Grevesse (1989); (b) Shale-normalized REE patterns. Normalization values for Post-Archean Australian Shale from Taylor and McLennan (1985); water depth listed after sample number in the Legend. (Color figure available online.)

using Ce anomalies calculated from bulk REE data as indicators of paleoredox conditions, given that the Ce anomaly is not only sensitive to variations in Ce content but also to variations in the concentrations of trivalent REEs, which are associated primarily with Fe and/or phosphate phases, whereas Ce is associated with the Mn phase. The Ce anomaly may not solely reflect Ce redox cycling and the REEs may provide indirect rather than direct evidence of changes in the depositional conditions that result from controls on the Fe/Mn ratio. Accordingly, the low REE contents of crust 3536-3 are most likely the result of hydrothermal or diagenetic contributions characterized by fast growth rates and high Ni and Li contents.

#### **Platinum-Group Elements**

Fe-Mn crusts are highly enriched in PGEs (Ir, Ru, Rh, Pt, Pd), especially Pt, compared to Earth's crustal abundances (Hein et al. 2000). Based on the five samples analyzed for PGEs, the Pt contents vary between 153 and 512 ppb, with an average value of 283 ppb; Rh, Ru, Pd, and Ir show concentrations up to 39, 21, 19, and 10 ppb, respectively (Table 4).

The processes of Pt enrichment in Fe-Mn crusts are not fully understood. Several mechanisms have been proposed for this enrichment, such as reduction or oxidation reactions, diagenetic or cosmogenic input, and enrichment related to phosphatization (Halbach et al. 1989a, 1990; Vonderhaar et al. 2000; Hein et al. 2005). We consider that the oxidative enrichment is the most likely general mechanism for high Pt concentrations; tetravalent Pt would be the final product, as is also the case for Ce, whereas Te is hexavalent (Hein et al. 2003). The most likely mechanism is that Pt is sorbed from seawater and then oxidized on the surface of the FeO(OH) (Hodge et al. 1985; Hein et al. 1997; Hein et al. 2003; Banakar et al. 2007).

#### **Interelement Relationships and Mineral Phases**

A correlation coefficient matrix for selected elements (Table 5) shows that Fe is positively correlated (99% Confidence Interval-CI) with Ti, Mo, and REEs (the correlation is better with LREEs and MREEs) and to a lesser extent (95% CI) with Y and Co. Manganese is positively correlated with W, Tl, and Pb at the 99% CI, and also with Zn, Mo, Ni, and Bi at 95% CI, and shows negative correlations with the REEs (except La, Pr, and Nd). Cobalt shows strong positive correlations (99% CI) with Te, Pb, Nb, Bi, and some REEs (La, Pr, Nd, Tb, Dy, Ho, Er, Yb), and the correlations are better with the MREEs and HREEs. Tellurium also shows positive correlations with the same elements as Co, but Te shows lower correlation coefficients with Bi and the REEs. Both Co and Te show higher coefficients with Fe than with Mn and both show negative correlations with elements characteristic of the detrital-aluminosilicate fraction, which was also found for Te in open-ocean crusts (Hein et al. 2003).

Q-Mode factor analysis produces five factors that explain 97.3% of total variance of the data. Considering the mineralogical composition of the samples, we conclude that the five factors represent the Fe oxyhydroxide, aluminosilicate, biogenic (which may also include hydrothermal Mn), hydrogenetic Mn, and CFA phases (Figure 6), which is in agreement with previous work (Muiños 2005) that used a slightly different dataset and statistical tool (Pereira et al. 2003) and with investigations of globally distributed crusts (Frank et al. 1999). It is also worth

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Table 4. PGE concentrations and ratios in five Fe-Mn crusts

$\Pr_{Au}$	57	
Pt/ Ru	11 24 12 23 23	1
Pd/ Ru	$   \begin{array}{r}     1.4 \\     0.67 \\     0.65 \\     0.56 \\     1.2   \end{array} $	1
Pt/ Ir	31 51 39 37 49	2
Pd/ Ir	3.8 1.4 1.7 2.5	1
Ru/ Rh	$1.0 \\ 0.54 \\ 0.68 \\ 1.0 \\ 0.48 \\ 0.48 \\ 0.48 \\ 0.28 \\ 0.48 \\ 0.48 \\ 0.28 \\ 0.$	
Pt/ Pd	8.1 37 21 22 22	1
Au (ppb)	6	
bd (dqq)	19 11 11 15	•
Pt (ppb)	153 512 231 223 223	ì
Rh (ppb)	14 39 18 18	ì
Ru (ppb)	14 21 17 18	•
Ir (ppb)	5 6 6	>
Fe/ Mn	$ \begin{array}{c} 1.04\\ 1.01\\ 1.62\\ 1.20\\ 1.35\\ \end{array} $	
Ce (wt%)	0.07 0.11 0.13 0.13 0.10	
Co (wt%)	0.26 0.41 0.20 0.17 0.43	2
Mn (wt%)	$ \begin{array}{c} 13.5 \\ 16.3 \\ 13.2 \\ 13.2 \\ 9.3 \\ 9.3 \\ \end{array} $	)
Fe (wt%)	14.0 16.6 21.4 15.7	
Depth (m)	1602 1839 2128 4594 4594	
Sample ID	3533-7 3478-B.3.5 3534-14.1 3513-14 3513-16	

I anto J.	CULLETE		N11100	TAT ITT	יז אווזא	11 orire	10 00	CITICITI													
	Fe	Mn	Fe/Mn	Si	AI	Ca	Ti	Ь	Bi	Co	Cu	Мо	ЧN	Ni	Pb	Te	Πh	Tl	M	Zn	Zr
Fe Mn	0.167												Correlation n= 18 for a	Coefficient all elements	matrix 18 si	amples					
Fe/Mn Si	0.720 - 0.449	-0.552 -0.567	0.023										95% confic 99% confic	dence level	s  0.468  s  0.582						
N	-0.473	-0.584	0.036	0.795																	
Ca	-0.434	0.058	-0.417	-0.426	-0.381																
Ti	0.648	-0.153	0.634	-0.234	-0.139	-0.316															
Р	-0.241	-0.023	-0.202	-0.358	-0.424	0.794	-0.145														
Bi	0.363	0.489	-0.030	-0.409	-0.459	-0.065	-0.064	0.108													
Co	0.503	0.362	0.187	-0.483	-0.434	-0.153	0.561	0.053	0.597												
Cu	-0.382	-0.291	-0.118	0.802	0.708	-0.549	-0.190	-0.339 -	-0.138 -	-0.283											
Mo	0.616	0.564	0.089	-0.313	-0.642	-0.323	0.075	-0.143	0.521	0.246 -	-0.148										
Nb	0.369	0.004	0.317	0.036	0.047	-0.442	0.664	-0.258	0.318	0.660	0.168	0.088									
Ņ	-0.540	0.519	-0.779	0.085	0.191	0.030	-0.423	-0.083	0.040	0.034	0.306 -	-0.131 0	.019								
Pb	0.442	0.587	-0.017	-0.567	-0.599	0.029	0.153	0.164	0.851	0.798 -	-0.356	0.441 0	.462	0.145							
Te	0.463	0.301	0.224	-0.402	-0.343	-0.181	0.469	0.000	0.547	0.928 -	-0.280	0.185 0	.692	0.122	0.809						
$\operatorname{Th}$	0.198	-0.458	0.512	0.320	0.337	-0.564	0.203	-0.418	0.084	0.135	0.395 -	-0.042 0	.225 -	0.302 -	-0.084	0.147					
Π	-0.018	0.722	-0.481	-0.205	-0.183	-0.233	-0.141	-0.088	0.486	0.465	0.165	0.369 0	.194	0.683	0.519	0.468	0.003				
W	0.411	0.821	-0.224	-0.546	-0.622	-0.057	-0.152	0.064	0.706	0.345 -	-0.294	0.738 0	.004	0.223	0.707	0.342 -	-0.280	0.626			
Zn	-0.005	0.572	-0.383	-0.051	-0.207	-0.116	-0.095	-0.078 -	-0.048 -	-0.004	0.140	0.363 0	.065	0.605	0.166	0.081 -	-0.471	0.477	0.427		
Zr	0.109	-0.289	0.315	0.027	0.122	-0.173	0.186	-0.014	0.399	0.406	0.242 -	-0.106 0	.346 –	0.166	0.289	0.357	0.679	0.146 -	-0.146	-0.395	
La	0.797	0.074	0.577	-0.561	-0.510	-0.184	0.721	0.002	0.404	0.642 -	-0.453	0.432 0	.413 –	0.542	0.471	0.525	0.276	-0.012	0.284	-0.315	0.275
Ce	0.757	-0.019	0.655	-0.029	-0.142	-0.722	0.506	-0.377	0.367	0.438	0.106	0.545 0	.390 -	0.412	0.282	0.400	0.652	0.183	0.285	-0.122	0.317
$\mathbf{Pr}$	0.771	0.011	0.605	-0.508	-0.418	-0.232	0.733	-0.034	0.362	0.614 -	-0.375	0.368 0	.392 -	0.559	0.403	0.486	0.383	-0.001	0.220	-0.377	0.347
PN	0.782	0.037	0.590	-0.527	-0.433	-0.225	0.728	-0.069	0.347	0.587 -	-0.405	0.385 0	.382 –	0.549	0.395	0.456	0.341	-0.037	0.230	-0.356	0.296
Sm	0.734	-0.079	0.624	-0.443	-0.318	-0.248	0.751	-0.119	0.265	0.504 -	-0.319	0.304 0	.374 –	0.609	0.263	0.347	0.394	-0.129	0.087	-0.435	0.363
Eu	0.569	-0.128	0.529	-0.285	-0.356	-0.110	0.684	0.078	0.328	0.542 -	-0.249	0.292 0	- 161	0.441	0.388	0.449	0.314	-0.093	0.094	-0.244	0.365
Gd	0.739	-0.050	0.621	-0.554	-0.363	-0.143	0.770	0.013	0.301	0.576 -	-0.417	0.230 0	.421 –	0.558	0.357	0.452	0.299	-0.119	0.131	-0.402	0.338
Tb	0.530	0.433	0.153	-0.665	-0.501	-0.067	0.324	0.198	0.520	0.622 -	-0.305	0.371 0	.116 –	0.045	0.590	0.510	0.211	0.499	0.576	-0.030	0.319
Dy	0.724	-0.040	0.607	-0.589	-0.392	-0.110	0.737	-0.010	0.315	0.611 -	-0.462	0.213 0	.394 –	0.545	0.384	0.492	0.330	-0.110	0.098	-0.430	0.399
Y	0.559	-0.053	0.476	-0.614	-0.390	0.124	0.693	0.234	0.338	0.649 -	-0.505	0.065 0	.440 –	0.435	0.433	0.531	0.113	-0.134	0.033	-0.427	0.371
Но	0.639	0.328	0.307	-0.653	-0.508	-0.103	0.676	0.142	0.217	0.611 -	-0.362	0.337 0	.273 -	0.173	0.402	0.478	0.114	0.317	0.371	0.085	0.166
Er	0.729	0.076	0.536	-0.607	-0.387	-0.109	0.728	0.051	0.420	0.682 -	-0.427	0.221 0	.408 –	0.444	0.474	0.524	0.205	-0.013	0.210	-0.347	0.341
Tm	0.493	0.002	0.366	-0.394	-0.335	0.027	0.669	0.076	0.200	0.423 -	-0.348	0.205 0	.504 -	0.292	0.254	0.325 -	-0.105	-0.227	0.018	-0.095	-0.022
$\mathbf{Y}\mathbf{b}$	0.706	-0.018	0.590	-0.599	-0.426	-0.021	0.743	0.114	0.358	0.627 -	-0.505	0.210 0	.471 –	0.548	0.434	0.526	0.239	-0.083	0.127	-0.415	0.395
Lu	0.293	-0.056	0.259	-0.349	-0.258	0.096	0.720	0.128	-0.032	0.486 -	-0.316	0.027 0	.521 -	0.163	0.118	0.403 -	-0.042	-0.058 -	-0.221	-0.082	0.142
H20 Depth	-0.002	-0.561	0.385	0.661	0.651	-0.626	0.093	-0.371	-0.167 -	-0.259	0.748 -	-0.146 0	.050 -	0.282 -	-0.455 -	0.323	0.658	-0.162 -	-0.371	-0.299	0.342

Table 5. Correlation Coefficient Matrix for selected elements

Yb Lu	0.654 0.141 -0.298 0.111 -0.298
Tm	0.667 0.791 0.791 0.791 0.791 0.791 0.791
Er	0.701 0.904 0.554 0.038 Fr
Но	0.760 0.434 0.723 0.723 0.723 0.723
Y	0.633 0.919 0.759 0.723 0.723 0.723
Dy	0.926 0.713 0.946 0.657 0.939 0.939 0.592 0.592
Tb	0.626 0.545 0.545 0.701 0.701 0.200 0.614 0.158 0.106
Gd	0.641 0.981 0.911 0.749 0.749 0.744 0.575 0.575
Eu	0.790 0.447 0.447 0.736 0.736 0.736 0.736 0.736 0.736 0.736 0.736 0.736
Sm	0.793 0.964 0.964 0.575 0.575 0.560 0.833 0.683 0.683 0.633 0.688 0.633 0.534 0.534 0.534 0.534 0.534 0.534 0.534 0.534 0.534 0.534 0.545 0.545 0.5550 0.5550 0.5550 0.5550 0.5550 0.5550 0.5550 0.5550 0.5550 0.5550 0.55500000000
PN	0.978 0.978 0.822 0.969 0.967 0.959 0.959 0.921 0.921 0.921 0.921 0.507 0.507
Pr	0.993 0.972 0.816 0.816 0.816 0.953 0.825 0.953 0.953 0.955 0.917 0.595 0.917 0.595 0.917
Ce	0.711 0.684 0.684 0.582 0.582 0.582 0.587 0.587 0.587 0.547 0.545 0.545 0.548 0.548 0.548 0.548 0.483
La	$\begin{array}{c} 0.672\\ 0.985\\ 0.989\\ 0.950\\ 0.957\\ 0.957\\ 0.957\\ 0.957\\ 0.957\\ 0.947\\ 0.848\\ 0.740\\ 0.913\\ 0.913\\ 0.530\\ 0.530\\ 0.530\\ 0.531\\ 0.642\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.642\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.664\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.662\\ 0.913\\ 0.913\\ 0.912\\ 0.$
Zr	$\begin{array}{c} 0.275\\ 0.317\\ 0.347\\ 0.347\\ 0.296\\ 0.365\\ 0.365\\ 0.365\\ 0.319\\ 0.365\\ 0.341\\ 0.341\\ 0.341\\ 0.341\\ 0.342\\ 0.341\\ 0.342\\ 0.$
	La La Pr Pr Nd Sm Eu Eu Dy Y Cd Dy Y Fr Tm Y tu UU H2O Depth

Continued
Table 5. (



Figure 6. Graphic display of Q-Mode rotated factor scores for five factors; CFA means carbonate fluorapatite.

noting that the fact that the number of samples is smaller than the number of variables does not affect the results, as discussed in Muiños (2005). The identified phases are comparable to the phases commonly found for Pacific and Indian Ocean Fe-Mn crusts.

#### Phosphatization

Phosphorus contents of our data set vary from 0.21 to 0.51 wt% (a factor of two) throughout the entire region and water depths (Figure 4). However, two samples, 3521-6 (Seine seamount) and 3533-7 (Annika seamount) are exceptions in that they show high P contents of 1.46 and 4.28 wt%, respectively, consistent with the CFA mineral component (Table 2). Regardless of these two samples, P seems to increase to the North and South of the Gibraltar: Samples from the MTR area (MTR-Josephine, Nameless, Unicorn, and Godzilla seamounts) show average P contents of 0.34 wt% while samples to the North (Tore) and South (Canaries) show higher average P values of 0.40 and 0.49 wt%, respectively.

#### S. B. Muiños et al.

Based on our limited dataset, phosphatization took place within the water-depth range of 1,200 to 1,500 m, but may also have occurred at shallower depths. That water-depth range corresponds to the lower core of MOW characterized by low dissolved oxygen contents and therefore low oxidation potential. This relationship was also pointed out by Koschinsky et al. (1996) and provides an indication of the influence of the prevailing oceanographic conditions in the area on the compositions of the crusts. Hein et al. (1993) suggested that during stable, warm climatic conditions dissolved phosphorous derived from intense chemical weathering on continents accumulated in the deep sea in large quantities. With the expansion of Antarctic glaciation and intensification of ocean circulation, the phosphorous-rich deep waters were redistributed by upwelling and turbulent mixing at the seamounts to intermediate water depths and may have been temporarily stored in the OMZ. Koschinsky and Halbach (1995) proposed a model for precipitation of hydrogenetic Fe-Mn crusts, with crusts being formed below the OMZ, as the result of the mixture of Mn<sup>2+</sup>-rich and O<sub>2</sub>-poor waters with  $Mn^{2+}$ -poor and O<sub>2</sub>-rich deep waters. Halbach et al. (1982, 1989b) and Koschinsky et al. (1997) linked the phosphatization with the expansion of the OMZ as the result of increased surface-water productivity. The expansion of the OMZ led to the impregnation with CFA of the extant crusts. Despite the fact that a phosphatized old crust generation is missing in Atlantic crusts, Koschinsky et al. (1996) noted phosphatization episodes within a 8.5 Ma record from a crust collected from Lion Seamount in the NE Atlantic Ocean, which may at least in part correspond to the 6 Ma event of phosphatization of limestone that occurred on Lighthill seamount off Morocco (Jones et al. 2002). The phosphatization was most likely a consequence of episodes of increased productivity and biogenic particle flux and is much younger than the phosphatization episodes in the Pacific (Hein et al. 1993).

Thus, the most probable explanation for shallow-water phosphatization is an interaction of the OMZ and MOW to produce an extended depth range for  $O_2$ -poor and  $Mn^{2+}$ -and dissolved P-rich waters, which reached down the slopes of some seamounts covered with Fe-Mn crusts. As a consequence, crust accretion may have been prohibited and precipitation of CFA promoted at these relatively shallow depths in the Atlantic Ocean (Hein et al. 2000 and references therein).

### **Resource Considerations**

Fe-Mn deposits in the Portuguese EEZ may become an important future resource but there is a clear need for studies to better understand their origin and distribution. During the oceanographic cruises mentioned above, as well as during the SO 83 cruise (Halbach and Scientific Crew 1993), samples of Fe-Mn nodules and crusts were collected on various seamounts from the northeast Atlantic. Despite the fact that the sampling was not systematic, and knowledge of the area needs to be augmented, we present a first-order evaluation of the possible resource potential of Fe-Mn crusts within and adjacent to the Portuguese EEZ, based on criteria developed by Hein et al. (2009).

#### Grade

According to Hein et al. (2009), large seamount summit areas with high grades of Co, Ti, REEs, Te, Ni, Th, Mn, Pt, etc, will be preferentially chosen for mining the crusts. It is also stressed that the grade will depend on the ability to collect Fe-Mn crusts without their substrate rocks, which would of course result in a decrease in the metal grade. Figure 7 shows the concentrations of selected trace metals that offer an economic potential for the future, for the crusts analyzed here compared to data for crusts from different ocean basins (from Hein 2004). Average metal concentrations in our dataset are comparable to average concentrations for Indian Ocean Fe-Mn crusts except for Zr, which is significantly higher in Indian Ocean crusts. For most of the metals, average concentrations for our dataset are in the range of the average Atlantic concentrations given by Hein (2004), except also for Zr and



**Figure 7.** Trace-metal concentrations in different ocean basins and in crusts analyzed here. Mean concentrations for Atlantic, Indian, and central Pacific crusts from Hein (2004). (Color figure available online.)

Pt. In comparison to central Pacific crusts, the average compositions of the crusts analyzed here are systematically lower, with the exceptions of Th, which is enriched in Atlantic and Indian crusts, and Te and Ce, which show similar concentrations. Nevertheless, if we consider the maximum values for our samples, the concentrations are similar to average central Pacific concentrations, except once again for Zr. The highest Co concentrations in our crusts of up to 0.54% Co are found in the shallower-water samples (on Seine and Nameless Seamounts; Figure 4). Further exploration is warranted within those areas and also regionally to possibly locate larger areas with mean crust grades similar to the maximum concentrations found for individual crusts analyzed here.

#### Tonnage (Thickness)

Tonnages were calculated based on average thicknesses of crusts in our dataset for Nameless, Unicorn, and MTR seamounts (Table 6). Considering data in Table 6, the surface area above 2,500 m water depth, and a mean wet-bulk density of crusts of  $1.95 \text{ g/cm}^3$ , we calculate the crust tonnage for Nameless, Unicorn, and MTR to be  $7.1 \times 10^7$ ,  $1.3 \times 10^8$ , and  $1.1 \times 10^9$  metric tons of wet crust respectively.

Considering the average concentrations for Co, Ni, Ce, Te, and Pt of 0.34%, 0.22%, 0.13%, 0.0043% and  $3 \times 10^{-80}$ %, respectively, we calculate maximum tonnages of these metals (Figure 8 and Table 6) with consideration for dilution resulting from recovery of substrate rock and unavailability of crusts for recovery due to other factors (see Hein et al. 2009 for discussion).

#### Area Permissible for Crust Coverage

Due to the absence of detailed sampling and the lack of backscatter side-scan sonar data, a comprehensive and more accurate calculation of the area of crust-coverage is not possible. Despite these limitations, we calculated surface areas using ArcMap's 3D analyst, ArcGIS<sup>®</sup> from ETOPO bathymetry and present those data considering reductions in areas potentially exploitable that result from water depth constraints, limitations in our knowledge of sediment cover and topography, and the necessity for biological corridors, based on the case study of Hein et al. (2009). The 2,500 m water depth limit proposed by Hein et al. (2009) is also used here even though some Fe-hosted metals increase with water depth, especially copper, which shows increased concentrations in crusts deeper than 3,000 m (Figure 4). We consider that the 2,500 m

Table 6.	Calculation of	of seamount a	nd ridge	surface	areas	for selec	cted sea	amoun	ts in
the study	area, crust m	ean thickness	es, and r	netric to	ns of s	selected	metals	based (	on a
crust me	an wet-bulk d	lensity of 1.95	$g/cm^3$						

Seamount	Mean Thickness (cm)	Surface area above 2,500 m water depth (km <sup>2</sup> )	Co (Tons)	Ni (Tons)	Ce (Tons)	Te (Tons)	Pt (Tons)
Nameless Unicorn	8 4	454 1948	$\begin{array}{c} 2.4\times10^5\\ 4.5\times10^5\end{array}$	$\begin{array}{c} 1.6\times10^5\\ 2.9\times10^7\end{array}$	$\begin{array}{c} 9.0\times10^6\\ 1.7\times10^5 \end{array}$	$\begin{array}{c} 3.0\times10^3\\ 5.7\times10^3\end{array}$	$\begin{array}{c} 2.1 \times 10^{-2} \\ 4.0 \times 10^{-2} \end{array}$
MTR	7	8492	$3.8 \times 10^6$	$2.5  imes 10^8$	$1.4 \times 10^6$	$4.8 \times 10^4$	$3.4 \times 10^{-1}$



Figure 8. Histogram of metric tons for different metals on the basis of average crust thicknesses for Nameless, Unicorn, and MTR seamounts.

water depth limit should be maintained because Cu is of limited economic potential either below or above the 2,500 m water depth and because the most economically important Fe-hosted metal -Te – does not show increasing concentrations with depth. In addition, most resource studies for the Pacific consider the 2,400–2,500 m isobath as the depth limit for calculations, and so comparisons of our data with Pacific data would not be valid using a different water depth limit.

We consider the following: (i) a large area represented by MTR (total surface area of 41,872 km<sup>2</sup>); (ii) a medium-sized area, represented by the combined areas of Lion, Nameless, Dragon, and Unicorn Seamounts (total surface area of 7,824 km<sup>2</sup>); and (iii) a small area represented by the combined areas of Lion and Nameless seamounts (total surface area of 3,676 km<sup>2</sup>) (Figure 9). The choice of Nameless, Lion, Dragon, and Unicorn seamounts was made because we expect that these seamounts located to the south of MTR may represent a suitable exploration area based on their depths, specific oceanic currents, and metal concentrations. Area calculations and reductions for the three scenarios are illustrated in Figure 10.

Hein et al. (2009) calculated surface areas for two case studies: (i) a large guyot (total surface area of  $11,761 \text{ km}^2$ ) and (ii) an average-size guyot (total surface area of  $3,495 \text{ km}^2$ ). Considering the worst-case scenario (50% reduction of crust-covered surface above 2,500 m water depth due to topographic and biological limitations for initial 60% sediment cover reduction), Hein et al. (2009) obtained, respectively,  $615 \text{ km}^2$  and



Figure 9. Madeira-Tore Rise; Lion and Dragon, and Nameless and Unicorn; and Lion and Nameless contour polygons used for surface area calculations. (Color figure available online.)



**Figure 10.** Histogram of total surface area and surface area above 2,500 m water depth for the three scenarios that consider different sizes of seamount areas. Other bars represent reductions in size of areas where crusts might be available because of sediment cover and other considerations (see Hein et al. 2009).

231 km<sup>2</sup> of permissible area. A worst-case scenario for our dataset would yield surface areas of 3,102, 1,116, and 540 km<sup>2</sup>, for the three groups described above, respectively (Figure 10). Considering these surface areas, a wet bulk density of  $1.95 \text{ g/cm}^3$ , an annual production of 1 million tons, and assuming a general mean crust thickness of 3 cm rather than using the limited dataset that we have for crust thicknesses from each seamount, the area needed to maintain a 20-year mine site is  $342 \text{ km}^2$ , which can be potentially accommodated by all three groups. These results are comparable to the results obtained by Hein et al. (2009) for central equatorial Pacific seamounts.

We also calculated metal tonnages considering the above conditions (worst-case scenario and 3 cm mean crust thickness). The metal concentrations used for these calculations are the ones given previously. We also calculated the dry-tonnages using a mean dry-bulk density of crusts of  $1.3 \text{ g/cm}^3$  (Hein et al. 2000). In Figure 11 we



Figure 11. Histogram of calculated tonnages (wet weight) for different metals considering the worst-case scenario of surface area reductions. Insets plots show areas corresponding to calculated tonnages based on dry weight.

present a histogram of calculated tonnages (wet- and dry-weight) for the metals considered. Maximum tonnages (resulting from large area calculations) for Co, Ni, Ce, Te and Pt are  $6.2 \times 10^5$ ,  $4.0 \times 10^5$ ,  $2.3 \times 10^5$ ,  $7.8 \times 10^3$  and  $5.4 \times 10^{-2}$  tons, respectively. Considering the small area, our calculations result in  $1.1 \times 10^5$ ,  $6.9 \times 10^4$ ,  $4.0 \times 10^4$ ,  $1.4 \times 10^3$  and  $9.5 \times 10^{-3}$  tons for the same metals, respectively. Our results show that the study area within the Portuguese EEZ (and adjacent areas in International Waters, for MTR; Figure 11) is comparable to that of areas in the central Pacific Ocean presented in Hein et al. (2009). Exploration beyond reconnaissance may now be warranted and should include detailed sampling, backscatter side-scan sonar, bathymetric mapping, and detailed mapping of crust thicknesses, etc., in order to better constrain the assumptions made here and to allow for a quantitative resource evaluation.

## Conclusions

The objective of this study was to determine the composition of Fe-Mn crusts from the northeast Atlantic and to consider gaps in our knowledge needed for assessing the quantitative resource potential of these deposits. The compositions of the studied crusts are typical for hydrogenetic crusts adjacent to continental margins. Specific compositional differences are found that likely indicate specific local conditions during crust accretion, for example higher Co, Ni, and Zr in Pacific crusts and higher Th in Atlantic crusts.

The enrichment of trace metals of economic interest in Fe-Mn crusts is of particular importance for their potential as a resource (i.e., Te and REEs), and is of specific interest for the resource potential of these deposits within the Portuguese EEZ. Based on the criteria of Hein et al. (2009), we calculated tonnages for specific metals in chosen areas in and adjacent to the Portuguese EEZ. Our results indicate that the study area is comparable to parts of the central Pacific Ocean and may represent an important metal resource for the future. Further studies are warranted in order to better constrain and quantify the results presented here.

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